

# HW #11

## 1. Bogoliubov Transformation

(a)

This is a straight-forward algebra

$$\begin{aligned}
 [b, b^\dagger] &= [a \cosh \eta + a^\dagger \sinh \eta, a^\dagger c \cosh \eta + a \sinh \eta] \\
 &= [a, a^\dagger] \cosh^2 \eta + [a^\dagger, a] \sinh^2 \eta \\
 &= \cosh^2 \eta - \sinh^2 \eta \\
 &= 1
 \end{aligned}$$

(b)

Using  $b, b^\dagger$  defined in (a), we can solve for  $a, a^\dagger$  as

$$a = b \cosh \eta - b^\dagger \sinh \eta$$

$$a^\dagger = b^\dagger \cosh \eta - b \sinh \eta$$

Substituting them into the Hamiltonian, we find

$$\begin{aligned}
 H &= \hbar \omega a^\dagger a + \frac{1}{2} V (a a + a^\dagger a^\dagger) \\
 &= \hbar \omega (b^\dagger \cosh \eta - b \sinh \eta) (b \cosh \eta - b^\dagger \sinh \eta) \\
 &+ \frac{1}{2} V ((b^\dagger \cosh \eta - b \sinh \eta)^2 + (b \cosh \eta - b^\dagger \sinh \eta)^2) \\
 &= b^\dagger b (\hbar \omega \cosh^2 \eta - V \cosh \eta \sinh \eta) \\
 &+ b b^\dagger (\hbar \omega \sinh^2 \eta - V \cosh \eta \sinh \eta) \\
 &+ b b (-\hbar \omega \sinh \eta \cosh \eta + \frac{1}{2} V (\sinh^2 \eta + \cosh^2 \eta)) \\
 &+ b^\dagger b^\dagger (-\hbar \omega \sinh \eta \cosh \eta + \frac{1}{2} V (\cosh^2 \eta + \sinh^2 \eta))
 \end{aligned}$$

Because  $\eta$  is a free parameter, we can choose it at our convenience. We choose  $\eta$  such that the coefficients of  $b b$  and  $b^\dagger b^\dagger$  vanish,

$$-\hbar \omega \sinh \eta \cosh \eta + \frac{1}{2} V (\cosh^2 \eta + \sinh^2 \eta) = 0.$$

$$-\hbar \omega \sinh 2\eta + V \cosh 2\eta = 0.$$

and hence

$$\eta = \frac{1}{2} \operatorname{arctanh} \frac{V}{\hbar \omega}, \text{ or } e^{4\eta} = \frac{\hbar \omega + V}{\hbar \omega - V}.$$

With this choice, the Hamiltonian has the familiar form

$$\begin{aligned}
 H &= b^\dagger b (\hbar \omega \cosh^2 \eta - V \cosh \eta \sinh \eta) \\
 &+ b b^\dagger (\hbar \omega \sinh^2 \eta - V \cosh \eta \sinh \eta) \\
 &= b^\dagger b (\hbar \omega (\cosh^2 \eta + \sinh^2 \eta) - 2 V \cosh \eta \sinh \eta) + (\hbar \omega \sinh^2 \eta - V \cosh \eta \sinh \eta) \\
 &= b^\dagger b \sqrt{(\hbar \omega)^2 - V^2} + \frac{1}{2} \left( \sqrt{(\hbar \omega)^2 - V^2} - \hbar \omega \right)
 \end{aligned}$$

We find that the ground state is annihilated by  $b$ , its energy is  $\frac{1}{2} \left( \sqrt{(\hbar \omega)^2 - V^2} - \hbar \omega \right)$ , and the excited states are obtained by acting  $b^\dagger$  each time raising the energy by  $\sqrt{(\hbar \omega)^2 - V^2}$ .

The system does not have a stable ground state if  $V > \hbar \omega$ .

(c)

We use the Hausdorff formula,

$$e^A B e^{-A} = B + [A, B] + \frac{1}{2!} [A, [A, B]] + \frac{1}{3!} [A, [A, [A, B]]] + \dots$$

In our case,  $A = (a a - a^\dagger a^\dagger) \eta / 2$ ,  $B = a$ .

Let us first calculate the multiple commutators.

$$B = a$$

$$[A, B] = \frac{\eta}{2} [a a - a^\dagger a^\dagger, a] = \eta a^\dagger.$$

$$[A, [A, B]] = \frac{\eta}{2} [a a - a^\dagger a^\dagger, \eta a^\dagger] = \eta^2 a.$$

$$[A, [A, [A, B]]] = \frac{\eta}{2} [a a - a^\dagger a^\dagger, \eta^2 a] = \eta^3 a^\dagger$$

etc. By inspection, it is clear that the  $2n$ -th commutator is  $\eta^{2n} a$ , while the  $2n + 1$ -th commutator is  $\eta^{2n+1} a^\dagger$ .

Now back to the Hausdorff formula,

$$e^A B e^{-A} = a + \eta a^\dagger + \frac{1}{2!} \eta^2 a + \frac{1}{3!} \eta^3 a^\dagger + \dots$$

$$= a \cosh \eta + a^\dagger \sinh \eta = b$$

Putting them together, we find  $b = U a U^{-1}$  indeed.

We can also see that it is a unitarity operator. The exponent is anti-Hermitian,

$$A^\dagger = \left( \frac{\eta}{2} (a a - a^\dagger a^\dagger) \right)^\dagger = \frac{\eta}{2} (a^\dagger a^\dagger - a a) = -\frac{\eta}{2} (a a - a^\dagger a^\dagger) = -A.$$

Therefore,  $U^\dagger = (e^A)^\dagger = e^{A^\dagger} = e^{-A} = U^{-1}$ , and hence it is unitary.

(d)

The ground state is annihilated by  $b$ , and hence  $U a U^{-1} | \text{gs} \rangle = 0$ .

Acting  $U^{-1}$  on both sides, we find  $a(U^{-1} | \text{gs} \rangle) = 0$ , namely that the state  $U^{-1} | \text{gs} \rangle$  is annihilated by the operator  $a$ . Therefore,  $U^{-1} | \text{gs} \rangle = | 0 \rangle$ ,

and hence

$$| \text{gs} \rangle = U | 0 \rangle.$$

Because  $U$  is a power series in  $\frac{\eta}{2} (a a - a^\dagger a^\dagger)$  which changes numbers by two units, we find that  $| \text{gs} \rangle$  can be expanded as

$$| \text{gs} \rangle = \sum_{n=0}^{\infty} c_{2n} | 2n \rangle$$

Using the fact that it is annihilated by  $b$ ,  $(a \cosh \eta + a^\dagger \sinh \eta) | \text{gs} \rangle = 0$ , we find

$$\begin{aligned} b | \text{gs} \rangle &= (a \cosh \eta + a^\dagger \sinh \eta) \sum_{n=0}^{\infty} c_{2n} | 2n \rangle \\ &= \sum_{n=0}^{\infty} c_{2n} (\sqrt{2n} | 2n-1 \rangle \cosh \eta + \sqrt{2n+1} | 2n+1 \rangle \sinh \eta) \\ &= \sum_{n=0}^{\infty} (c_{2n+2} \sqrt{2n+2} \cosh \eta + c_{2n} \sqrt{2n+1} \sinh \eta) | 2n+1 \rangle = 0. \end{aligned}$$

It means

$$c_{2n+2} = c_{2n} \left( -\sqrt{\frac{2n+1}{2n+2}} \tanh \eta \right).$$

Therefore,

$$c_{2n} = c_0 (-1)^n \sqrt{\frac{(2n-1)!!}{(2n)!!}} \tanh^n \eta$$

The norm is

$$1 = c_0^2 \sum_{n=0}^{\infty} \frac{(2n-1)!!}{(2n)!!} \tanh^{2n} \eta$$

Rewriting it with the Gamma function,

$$c_0^{-2} = \sum_{n=0}^{\infty} \frac{2^n \Gamma(n+1/2) / \sqrt{\pi}}{2^n \Gamma(n+1)} \tanh^{2n} \eta = \frac{1}{\sqrt{\pi}} \sum_{n=0}^{\infty} \frac{\Gamma(n+1/2)}{\Gamma(n+1)} \tanh^{2n} \eta$$

*Mathematica* gives this infinite series

$$\text{In}[28] := \text{Sum} \left[ \frac{\text{Gamma} \left[ n + \frac{1}{2} \right]}{\text{Gamma} [n + 1]} x^n, \{n, 0, \infty\} \right]$$

$$\text{Out}[28] = \frac{\sqrt{\pi}}{\sqrt{1-x}}$$

Therefore,

$$c_0^{-2} = \frac{1}{\sqrt{\pi}} \frac{\sqrt{\pi}}{\sqrt{1-\tanh^2 \eta}} = (1 - \tanh^2 \eta)^{-1/2}$$

and hence

$$c_0 = (1 - \tanh^2 \eta)^{1/4}.$$

Putting everything together, the ground state of the Hamiltonian is given in terms of the number states

$$| \text{gs} \rangle = (1 - \tanh^2 \eta)^{1/4} \sum_{n=0}^{\infty} (-1)^n \sqrt{\frac{(2n-1)!!}{(2n)!!}} \tanh^n \eta | 2n \rangle.$$

The Bogoliubov transformation appears not only in the BEC but also in optics. The laser is a coherent state, and one can realize Bogoliubov transformed ones called the "squeezed states." They have been under active research.

## 2. Superfluid Vortex

(a)

The solution is  $z$ -independent, and hence is a function of  $r$  and  $\theta$  only. We use the chain rule

$$\begin{pmatrix} \nabla_r \\ \nabla_\theta \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -r \sin \theta & r \cos \theta \end{pmatrix} \begin{pmatrix} \nabla_x \\ \nabla_y \end{pmatrix}$$

and find

$$\begin{pmatrix} \nabla_x \\ \nabla_y \end{pmatrix} = \frac{1}{r} \begin{pmatrix} r \cos \theta & -\sin \theta \\ r \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nabla_r \\ \nabla_\theta \end{pmatrix}.$$

Therefore,

$$\begin{pmatrix} \nabla_x \\ \nabla_y \end{pmatrix} \psi = \begin{pmatrix} \nabla_x \\ \nabla_y \end{pmatrix} f(r) e^{in\theta} = \frac{1}{r} \begin{pmatrix} r \cos \theta & -\sin \theta \\ r \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} f'(r) \\ i n f(r) \end{pmatrix} e^{in\theta}.$$

The momentum density is

$$\vec{j} = \frac{\hbar}{2mi} f(r) \frac{1}{r} \begin{pmatrix} r \cos \theta & -\sin \theta \\ r \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} f'(r) \\ i n f(r) \end{pmatrix} + c.c.$$

$$= \frac{n\hbar}{mr} f(r)^2 \begin{pmatrix} -\sin \theta \\ \cos \theta \end{pmatrix}.$$

On the other hand, the number density is

$$\rho = \psi^* \psi = f(r)^2.$$

The velocity field is their ratio,

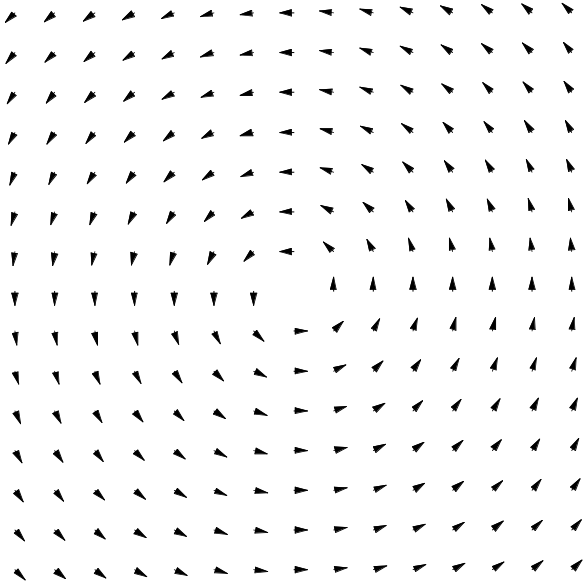
$$\vec{v} = \frac{\vec{j}}{\rho} = \frac{n\hbar}{mr} \begin{pmatrix} -\sin \theta \\ \cos \theta \end{pmatrix}.$$

For the plotting purpose, we take  $\hbar = m = 1$ .

`In[7] := << Graphics`PlotField``

`n = 1`

```
In[8]:= PlotVectorField[ $\frac{1}{\sqrt{x^2 + y^2}}$  {-y, x}, {x, -1, 1},  
  {y, -1, 1}, ScaleFactor -> None, ScaleFunction -> (.05 # &)]  
  
Power::infy : Infinite expression  $\frac{1}{0}$  encountered. More...  
  
∞::indet : Indeterminate expression 0 ComplexInfinity encountered. More...  
  
∞::indet : Indeterminate expression 0 ComplexInfinity encountered. More...
```



```
Out[8]= - Graphics -
```

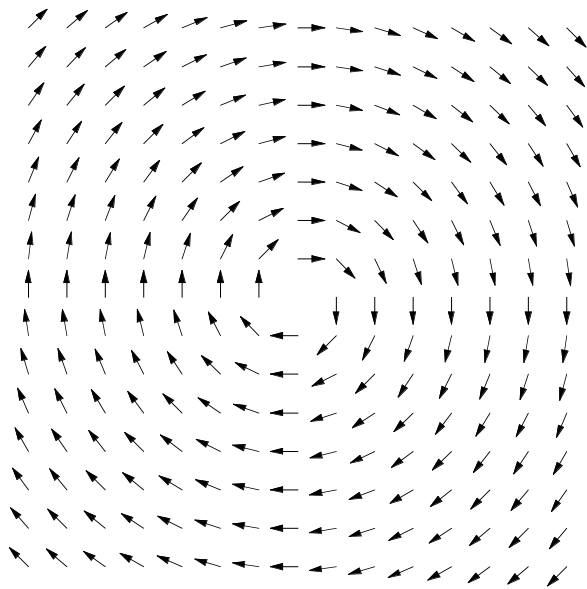
```
n = -2
```

```
In[9]:= PlotVectorField[ $\frac{-2}{\sqrt{x^2 + y^2}}$  {-y, x}, {x, -1, 1},  
  {y, -1, 1}, ScaleFactor -> None, ScaleFunction -> (.05 # &)]
```

```
Power::infty : Infinite expression  $\frac{1}{0}$  encountered. More...
```

```
∞::indet : Indeterminate expression 0 ComplexInfinity encountered. More...
```

```
∞::indet : Indeterminate expression 0 ComplexInfinity encountered. More...
```



```
Out[9]= - Graphics -
```

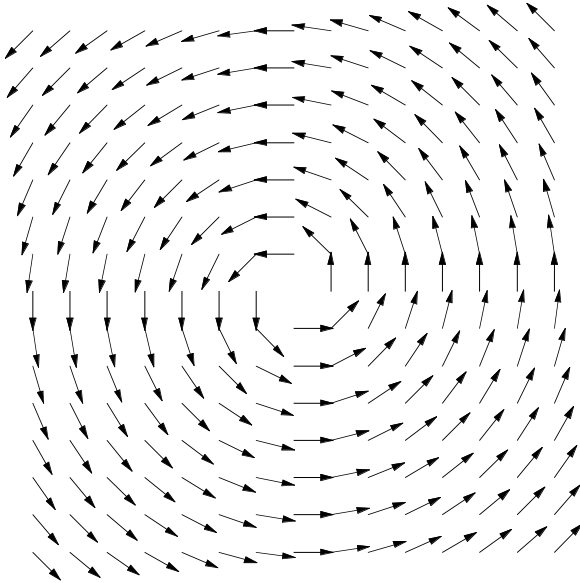
```
n = 3
```

```
In[10]:= PlotVectorField[ $\frac{3}{\sqrt{x^2 + y^2}}$  {-y, x}, {x, -1, 1},
  {y, -1, 1}, ScaleFactor -> None, ScaleFunction -> (.05 # &)]
```

Power::infy : Infinite expression  $\frac{1}{0}$  encountered. More...

∞::indet : Indeterminate expression 0 ComplexInfinity encountered. More...

∞::indet : Indeterminate expression 0 ComplexInfinity encountered. More...



Out[10]= - Graphics -

(b)

The circulation is obtained from the loop integral of the velocity field

$$\vec{v} = \frac{\vec{j}}{\rho} = \frac{n\hbar}{mr} \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix},$$

$$\kappa = \oint \frac{n\hbar}{mr} \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} \cdot r \begin{pmatrix} -\sin\theta \\ \cos\theta \end{pmatrix} d\theta$$

$$= \frac{2\pi n\hbar}{m} = \frac{n\hbar}{m}$$

(c)

We start with the Euler-Lagrange equation obtained from the Lagrangian,

$$i \hbar \dot{\psi} + \frac{\hbar^2 \Delta}{2m} \psi + \mu \psi - \lambda \psi^* \psi \psi = 0.$$

We now substitute our Ansatz into the equation:  $\psi(r, \theta, t) = e^{in\theta} f(r)$ .

Because it is static  $\dot{\psi} = 0$ , we find

$$\frac{\hbar^2}{2m} \left( \frac{d^2}{dr^2} + \frac{1}{r} \frac{d}{dr} - \frac{n^2}{r^2} \right) f + \mu f - \lambda f^3 = 0.$$

We now try to cast it to a form as independent of the parameters as possible.

$$\left( \frac{d^2}{dr^2} + \frac{1}{r} \frac{d}{dr} - \frac{n^2}{r^2} \right) f + \frac{2m\mu}{\hbar^2} f - \frac{2m\lambda}{\hbar^2} f^3 = 0$$

We introduce the variable

$$\frac{\sqrt{2m\mu}}{\hbar} r = \rho$$

Then the equation becomes

$$\left( \frac{d^2}{d\rho^2} + \frac{1}{\rho} \frac{d}{d\rho} - \frac{n^2}{\rho^2} \right) f + f - \frac{\lambda}{\mu} f^3 = 0$$

Finally, we introduce the function

$$f = \sqrt{\frac{\mu}{\lambda}} \phi$$

which satisfies

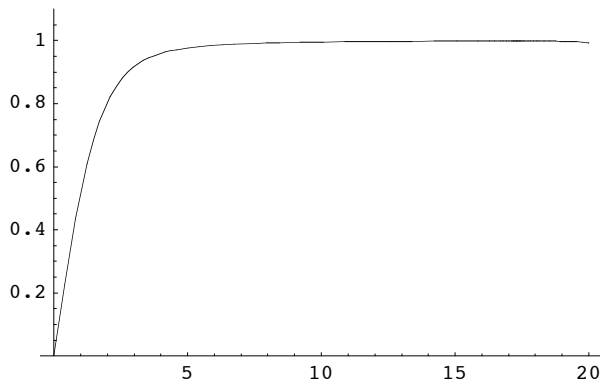
$$\left( \frac{d^2}{d\rho^2} + \frac{1}{\rho} \frac{d}{d\rho} - \frac{n^2}{\rho^2} \right) \phi + \phi - \phi^3 = 0$$

The equation is now completely free from physical parameters except for the integer  $n$ .

We solve numerically with the boundary conditions  $\phi(0) = 0$ ,  $\phi(\infty) = 1$ . Instead of imposing the boundary condition strictly at the origin, which causes all kinds of problems, we cheat a little bit and impose it slightly off the origin.

```
vortex = NDSolve [ { y ' ' [ x ] +  $\frac{1}{\mathbf{x}}$  y ' [ x ] -  $\frac{1}{\mathbf{x}^2}$  y [ x ] + y [ x ] - y [ x ]3 == 0 ,
  y [ 10-6 ] == 0 , y ' [ 10-6 ] == 1.166380374636 } , y , { x , 10-6 , 20 } ]
{ { y → InterpolatingFunction[ { { 1. × 10-6 , 20. } } , <> ] } }

Plot [ y [ x ] /. vortex [ [ 1 ] ] , { x , 10-6 , 20 } , PlotRange → { 0 , 1.1 } ]
```



- Graphics -

This way, we see that the vortex solution exists for the classical Schrödinger field with a repulsive self-interaction potential, as observed in the atomic BEC as well as liquid He-II.